

***** 2026 May 28 Thursday Night Net Solar Report ***** [K9JU]

Unless otherwise noted, the following information comes from <https://solarham.com/>. The Hp30 index comes from the GFZ, a geophysical group in Germany. Other sources may include <https://www.wm7d.net/>, <https://spaceweather.com/> and <https://www.spaceweatherlive.com/>. See those websites for further details. All times are EDT (UTC - 4).

- The Solar Flux Index (SFI) ¹ at 1600 ET (2000 UTC) is 145 (+1 from yesterday).
- The Wolf (or Sunspot or Zurich) Number ⁽²⁾ was 165 (+17 from yesterday (+98 from last Thursday) from eleven Active Regions (down one from yesterday) at the time of publishing the number (8 at 1800). Number of individual sunspots = Wolf number - [10 * (number of ARs)] or, 165 - [10 * (11)] = 55.
- X-ray flare activity remains low for the most part with flares in the C-class riding on on a high B-class pedestal. However, at 1420 ET, there was an almost “M-class” flare, measuring C9.6). (NOTE: A-, B-, and C-class flares do not cause Radio Blackouts (RBO). but may add to the SFI. For further developments, keep an eye on <https://solarham.com/>, <https://spaceweather.com/> and, for much more up to date solar conditions, visit [://www.spaceweatherlive.com/](https://www.spaceweatherlive.com/) (click “Solar activity” and then “Real-time Solar activity” on the drop down menu).
- At 1830 ET, the geomagnetic field was active with a K_p ⁽³⁾ = 3.00. The K_p has a 3 hour measurement period. The Hp30 ⁽²⁾ index was also indicating active conditions with a reading of 3.67. The Hp30 has a 30 minute measurement period.
- The solar wind ⁽⁹⁾ parameters at 1830: the wind speed was elevated above the average background speed at 450 km/sec (up from 380 km/sec this morning (the background average is ~300 km/s); the density was near 4.7 p/cm³ or just over one and a third times the long term average of ~3.14 p/cm. The B_z component of the Interplanetary Magnetic Field ¹⁶ was negative (south pointing) at -4nT after being predominantly south pointing for the preceding 3-hours. For an explanation of what parameter values might lead to geomagnetic storms, see footnote (*) below.
- There still are three Coronal Holes ⁽⁷⁾ (#58, #60 and #61) on the face of the sun. One (#60) is located just east of the north/south center line of the sun and should be centered in a day or two; this CH is in the northern hemisphere of the sun. The second (#58) is long and looked like a smoking pipe and located slightly west of the central north/south line of the sun and in the southern hemisphere of the sun. The third (#61) is located on the equator of the sun and past the halfway mark between central sun and the eastern limb of the sun. Due to locations, the first two CHs shouldn't cause too much trouble for earth, only a bit of side splash. The third may effect earth in 2-3 days as it moves through central sun although the CH is rather small. There remain several filaments showing on the face of the sun but none has been reported as leaving the surface of the sun.
- Summary: With the number of sunspots and SFI hanging at relatively good levels, radio propagation conditions for daytime DX should be fairly decent on the higher HF bands (20M and 10M). 10M and 12M should be north to south propagation. 20M and 17M will hold the bulk of intercontinental communications propagation. Nighttime and early morning DX

should be good on 40M and 30M; DX on 160M and 80M conditions should be limited due to increased band noise from thunderstorms.

As the maximum of Solar Cycle 25 has occurred, conditions will, for the most part, range from good to very good for another year or so especially those months between the Autumnal Equinox (09/22), through the winter solstice and until mid to late May next year (***)). Enjoy these HF DX conditions as they won't last forever! The solar minimum is forecast to be in 2030, just over 4-years away.

For up to date info, stay tuned to <https://www.solarham.net/>, <https://spaceweather.com/> and <https://www.spaceweatherlive.com/>

¹ Solar Flux Index (SFI): See https://glossary.ametsoc.org/wiki/Solar_flux

² Wolf [Sunspot] Number: See https://en.wikipedia.org/wiki/Wolf_number and <https://ngdc.noaa.gov/stp/solar/ssndata.html>. From brief research on the time of day the number is published, it appears the time is 1230 UTC (8:30 AM ET). The number also appears to be based upon multiple observations throughout a 24-hours day.

³ K_p is the planetary K-index. See <https://www.swpc.noaa.gov/products/planetary-k-index> and <https://seetheaurora.com/kp-index-explained>.

⁴ Solar flare definition and classes at <https://www.spaceweatherlive.com/en/help/what-are-solar-flares.html#:~:text=A solar flare is defined as a sudden%2C,up in the solar atmosphere is suddenly released. Also, read https://www.space.com/solar-flares-effects-classification-formation>.

⁵ Radio Black Out categories may be found at the SMARC website under “Files” and “Presentations.”

⁶ Geomagnetic Storm categories may be found at the SMARC website under “Files” and “Presentations.”

⁷ Coronal Holes (CH): See <https://www.swpc.noaa.gov/phenomena/coronal-holes>

⁸ Tamitha Skov, WX6SWW, <https://www.spaceweatherwoman.com/>

⁹ Solar wind parameters at <https://www.solarham.net/solarwind.htm> and explanations of the parameters at <https://hypertextbook.com/facts/2005/RandyAbbas.shtml>

¹⁰ CME (Coronal Mass Ejection): See: https://en.wikipedia.org/wiki/Coronal_mass_ejection

¹¹ Carrington Event: See https://en.wikipedia.org/wiki/Carrington_Event. YouTube video explaining in simple terms the Carrington Event and solar storms: <https://www.youtube.com/watch?v=85-p9EIEVUA> [thank you to Joe, WB9VVH - ed].

¹² Hp30 and Hp60 Indexes: See <https://kp.gfz-potsdam.de/en/hp30-hp60>

¹³ CIR (Co-rotational Interactive Region): See [https://www.swpc.noaa.gov/news/coronal-hole-high-speed-streams-ch-hss#:~:text=Persistent coronal holes are long-lasting sources for high,forms%2C known as a co-rotating interaction region \(CIR\)](https://www.swpc.noaa.gov/news/coronal-hole-high-speed-streams-ch-hss#:~:text=Persistent%20coronal%20holes%20are%20long-lasting%20sources%20for%20high-speed%20streams%20known%20as%20a%20co-rotating%20interaction%20region%20(CIR))

¹⁴ Solar Dynamics Observatory Helioseismic and Magnetic Imager (SDO HMI): https://en.wikipedia.org/wiki/Solar_Dynamics_Observatory

¹⁵ Magnetic Filaments: <https://spaceweather.com/glossary/filaments.html>.

¹⁶ Interplanetary Magnetic Field (IMF): <https://www.spaceweatherlive.com/en/help/the-interplanetary-magnetic-field-imf.html>

¹⁷ Solar filaments or prominences: https://en.wikipedia.org/wiki/Solar_prominence

¹⁸ SWPC: Space Weather Prediction Center: <https://www.swpc.noaa.gov/>.

¹⁹ Parker Spiral: https://en.wikipedia.org/wiki/Heliospheric_current_sheet.

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(*) Thermally driven high velocity thermospheric winds at F2 region altitude blow from the high temperature dayside hemisphere to the lower temperature nightside hemisphere reducing the volume of neutral atomic oxygen available to be ionized in the F2 region during daytime hours, lowering the MUF for F2 propagation crossing the northern hemisphere to mostly less than 24 MHz through mid-September and mostly less than 28 MHz through late September.

(**) Sustained southward orientation (-Bz) of the north-south component of the interplanetary magnetic field (IMF) plays a crucial but unpredictable role in triggering all geomagnetic storms. Brief minor to moderate geomagnetic storms may be gradually triggered when the IMF is sustained in a southward orientation (-Bz) with an IMF field strength of about -5 nanoteslas or less for at least a few hours coincident with the influence of a geoeffective coronal hole high speed stream or CME. More frequent, longer duration, minor to severe geomagnetic storms may be triggered suddenly and unpredictably when the IMF is sustained in a southward orientation (-Bz) with IMF field strength of about 5 nanoteslas or more for a period of several hours or more coincident with the influence of a geoeffective CME and solar wind speed of about 500 km/second or more.

(***) Comment and information belong to Frank Donovan, W3LPL, and all credit for the content is his.

For an explanation of the processes involved when a CME hits the magnetic field of the earth, watch <https://www.youtube.com/watch?v=j59pEqkPkLE>. Thanks to Joe, WB9VVH, for the link.

***** DX *****

The following information is from several sources including DX World <https://www.dx-world.net/>, DXNews <https://dxnews.com/>, the ARRL DX Bulletin <http://www.arrl.org/w1aw-bulletins-archive-dx> and/or personal experience. Most, if not all, DXpeditions will operate SPLIT; refer to the operating manual for your radio for instructions. Do NOT call on the frequency of the DXpedition as they are not listening on their own frequency! NOTE: Following the call of the DXpedition, in parenthesis, are the principal prefix, the general location plus latitude/longitude, beam heading (°) from East TN, and the

ClubLog Most Wanted list standing out of 340 for the country/entity. The Most wanted, #1, is P5, North Korea and least needed, #340, is the USA.

- N5J, Jarvis Islands is over. For QSL more information, see <https://jarvisisland2024.com/qsl-info-log-search/log-oqrs>.
- CY9C, St. Paul Island (CY9, island near Nova Scotia, Canada at 47N/60W, 50°, #50), 08/26 until 09/05. Operations 160M to 6M using CW, SSB and FT8 ¹. For more information, see <https://t-rexsoftware.com/cy9c/index.htm>.

¹ F/H (Fox and Hound or DXpedition) mode. See <https://www.onallbands.com/using-the-ft8-dxpedition-mode-a-k-a-fox-and-hound/>. Keep in mind, you are the hound.

² MSHV (Multi Stream HV) mode. See <http://lz2hv.org/node/10>.

³ Find a PDF user guide in the Version 2.7.0-rc6 section <https://wsjt.sourceforge.io/wsjt.html>. Scroll down to the “Candidate Release: WSJT-X 2.7.0-rc6” section.

⁴ Download the Win 64-bit software at <https://sourceforge.net/projects/wsjt/files/wsjt-2.7.0-rc6/wsjt-2.7.0-rc6-win64.exe/download>. Versions for Mac O/S and Linux are also available.

For a bar graph of DXpeditions, on a per month basis, go to <https://www.dx-world.net/>.

Remember: Technician Class Amateur Radio Operators have (somewhat limited) operating privileges to use phone (**only** SSB) and digital modes on the 10M band using up to 200-Watts PEP output power. Your segments are digital from 28.000-MHz to 28.300-MHz (stay near the digital mode frequencies above 28.070-MHz out of courtesy to CW operations) and SSB from 28.300-MHz to 28.500-MHz. You are not authorized to use FM mode as it is above your phone cutoff frequency. 10M has experienced exceptional propagation to all parts of the world during day light hours. The conditions should continue for another 2- to 3-years. A simple 10M dipole is only 16’ in length and can be fed in the center with cheap RG-6/U. “But it ain’t 50-Ohms coax?” So what! The SWR is only 1.5:1 which any HF radio should handle. The loss characteristics are better than RG-8X. Get on 10 and have fun! You just might get the DX bug and upgrade to General Class. A win-win situation for you.

For more DXpeditions, from DX World, <https://www.dx-world.net/>, please see [http://www.hamradiotimeline.com/timeline/dxwSouth Pacific Ocean at timeline 1 1.php](http://www.hamradiotimeline.com/timeline/dxwSouth%20Pacific%20Ocean%20at%20timeline%201%201.php).

***** **ARRL Letter** *****

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- The ARRL podcast, “On The Air,” is new with Joe, NJ1Q, W1AW Station Manager, tells us ways to “Practice Code with W1AW” code practice sessions. The podcasts with iTunes for iOS, Stitcher for Android and blubrry for desktop computers.]

- New editions of the ARRL Audio News,” a summary of the top news for the week of ham radio and the ARRL, appear every Friday. All editions of the Audio News are available at the blubrry website, <https://blubrry.com/arrlaudionews/>. The latest Audio News may be heard locally on the W4KEV repeater (Knoxville), 145.370-MHz, on Fridays, Saturdays and Sundays at 8:30 AM and 10:00 PM ET.
- ARRL has now expanded access to the publications archives to members of the League. Both QEX (technical - 1981 to 2011) and NCJ (National Contest Journal) - 1973 to 2011) are available under “ARRL Periodicals Archive and Search.”
- Processing of DXCC award applications, whether paper, online or LoTW, is still stalled. In the last communications from the League, there is no known date for return to processing DXCC award applications. In the meantime, you can upload your contacts to LoTW. This has to do with the method of combining paper (QSL cards), online DXCC and LoTW applications into the Master DXCC listings. The process does not effect WAS and VUCC applications and certificate issuances.

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